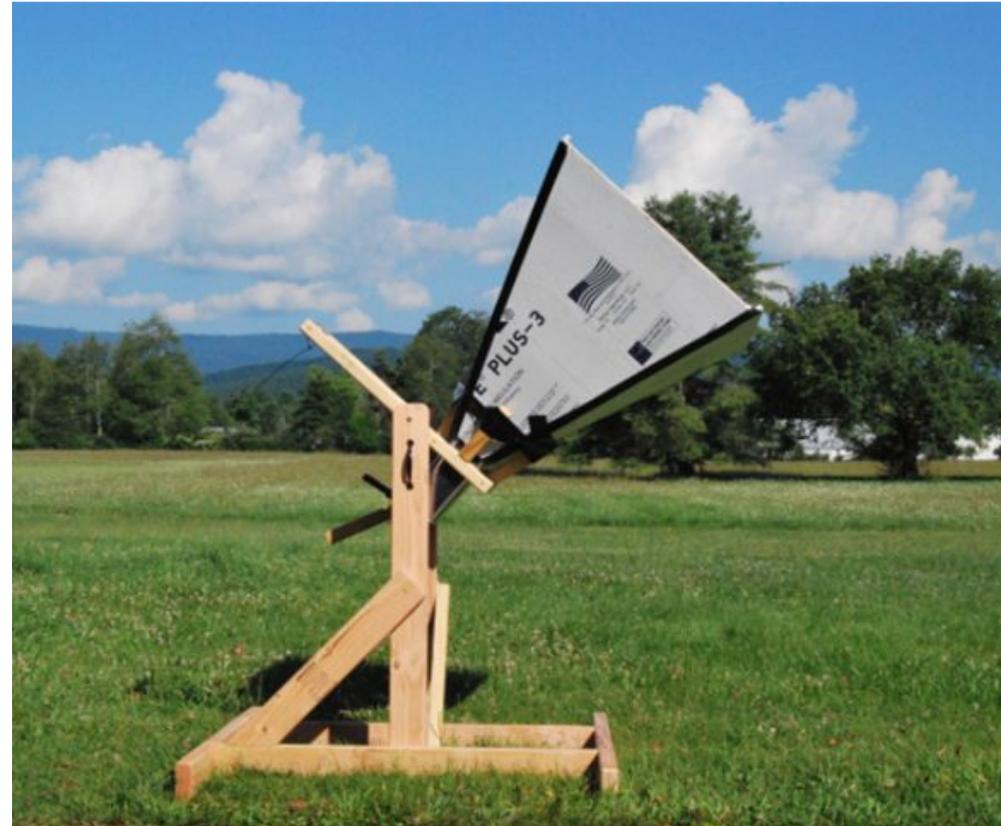
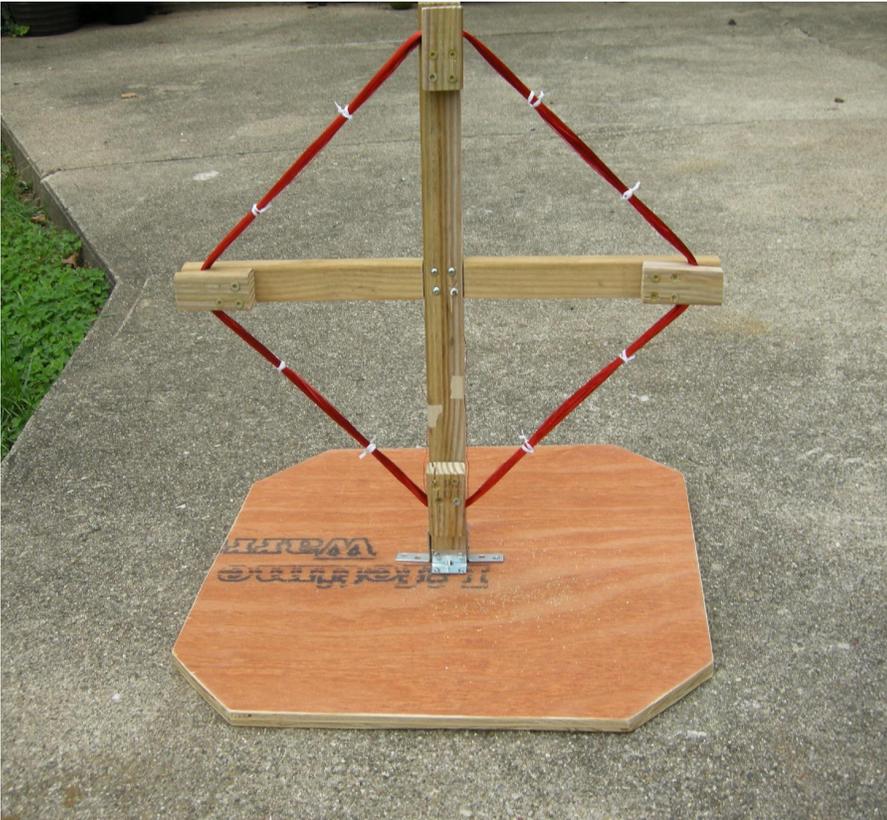


*observe the invisible universe...*

# OPEN SOURCE RADIO TELESCOPES



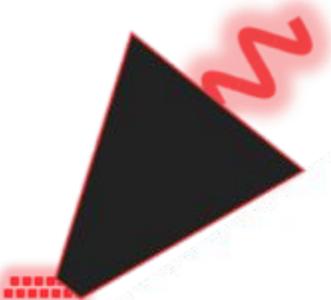
**Evan Smith, Ellie White, Richard Prestage, & Martin Braun  
FOSDEM 2018**



# What is Open Source Radio Telescopes?

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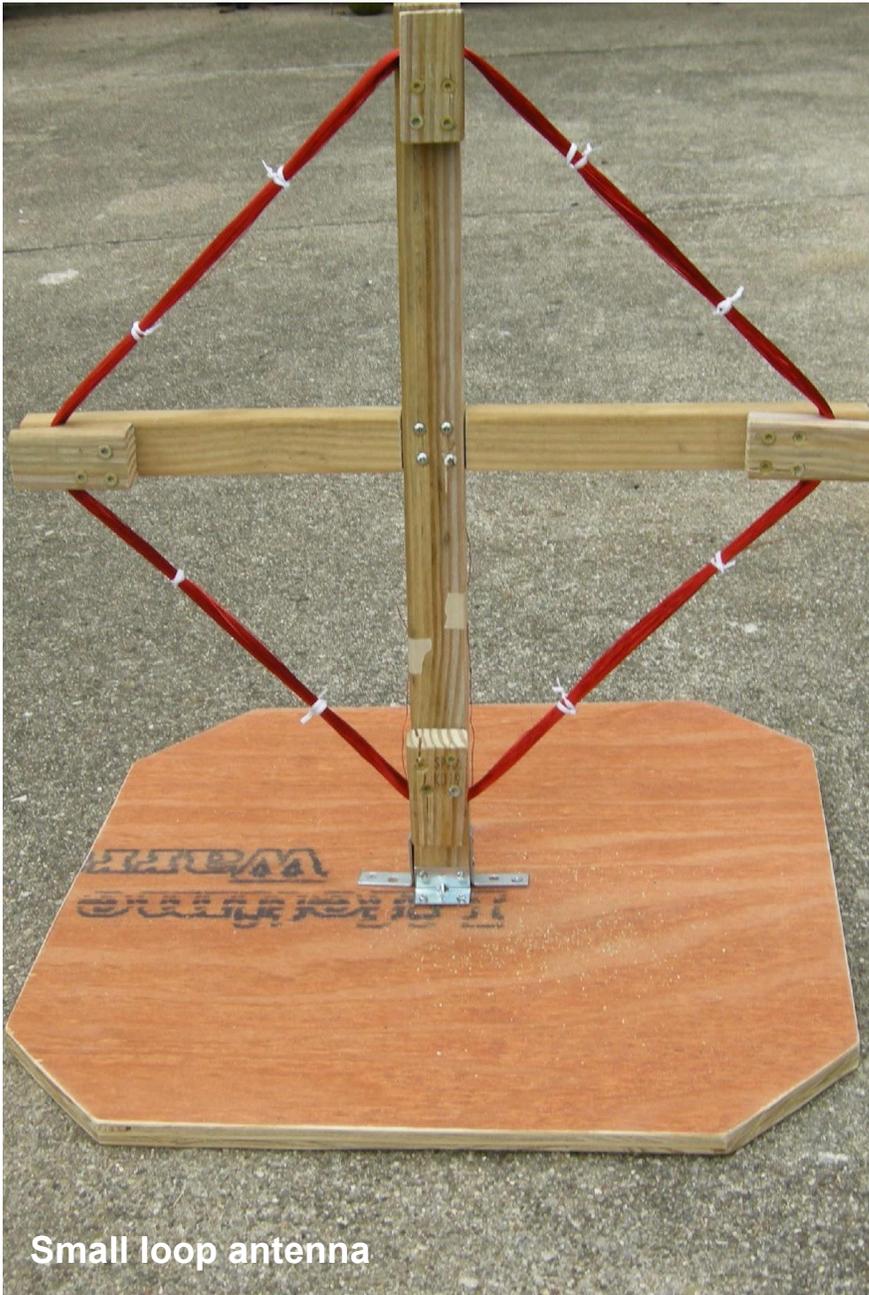
- An open, collaborative, and free collection of ideas and designs for radio telescope construction
- OSRT provides plans and learning resources that are accessible to anyone with an interest in citizen science or STEM education, from middle school to graduate students and beyond
- OSRT promotes discussions about digital signal processing and using GNU Radio software with off-the-shelf electronics, such as low-noise amplifiers, filters, and software defined radios.



## Some projects we are currently pursuing...

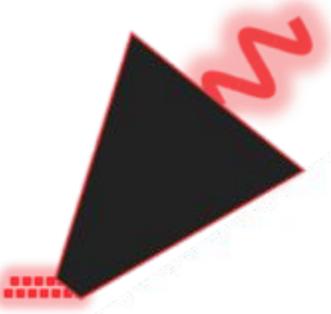
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- Designed, built, and tested a 21cm Horn Antenna, capable of detecting neutral hydrogen and mapping the Milky Way
- Developed a Small Loop Antenna, which is capable of detecting solar activity by monitoring the signal strength of Very Low Frequency (VLF) submarine stations.
- In the process of developing instruction manuals and kit prototypes for both antenna designs, which we will distribute to students and teachers to promote STEM education.
- Each of these projects involves the use of RTL-SDR dongles and GNU Radio flowgraphs.



# The Small Loop Antenna

(< \$100, easy to make, middle-school level)



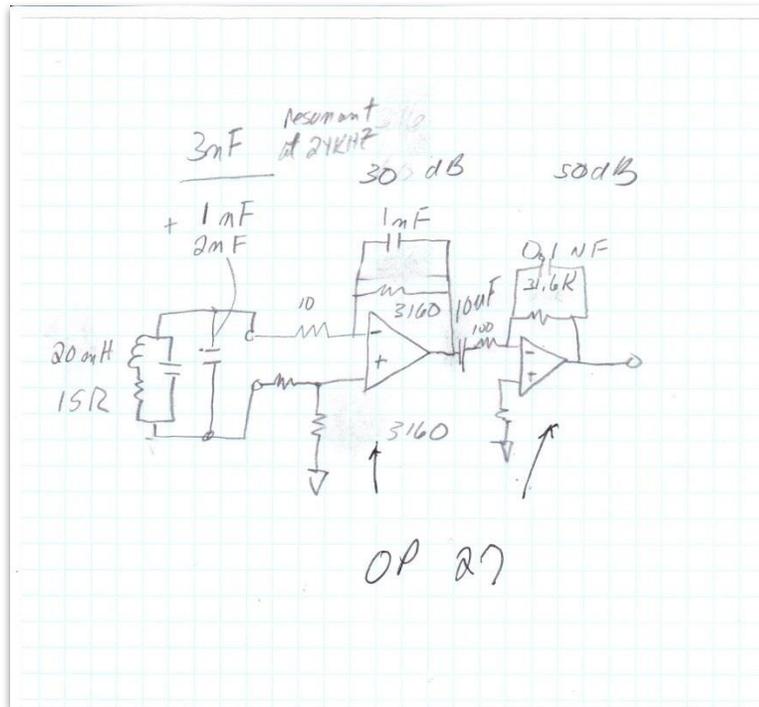
# Physical construction of the loop

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- Cross-frame built with scrap wood
- Frame wound with 125 turns of 24 awg magnet wire
- Mounted antenna to baseboard with L-brackets

# Tuning/Amplifying circuit



Tuning/amplifying circuit schematic (credit Steve White, Green Bank Observatory)

- Steve White of Green Bank Observatory built the tuning-amplifying circuit for the antenna.
- The LC circuit is tuned to 24 kHz, the frequency of the 2MW Cutler, Maine VLF station.
- Op-amps were used to amplify the signal enough for us to detect it.

# How the loop antenna detects solar flares

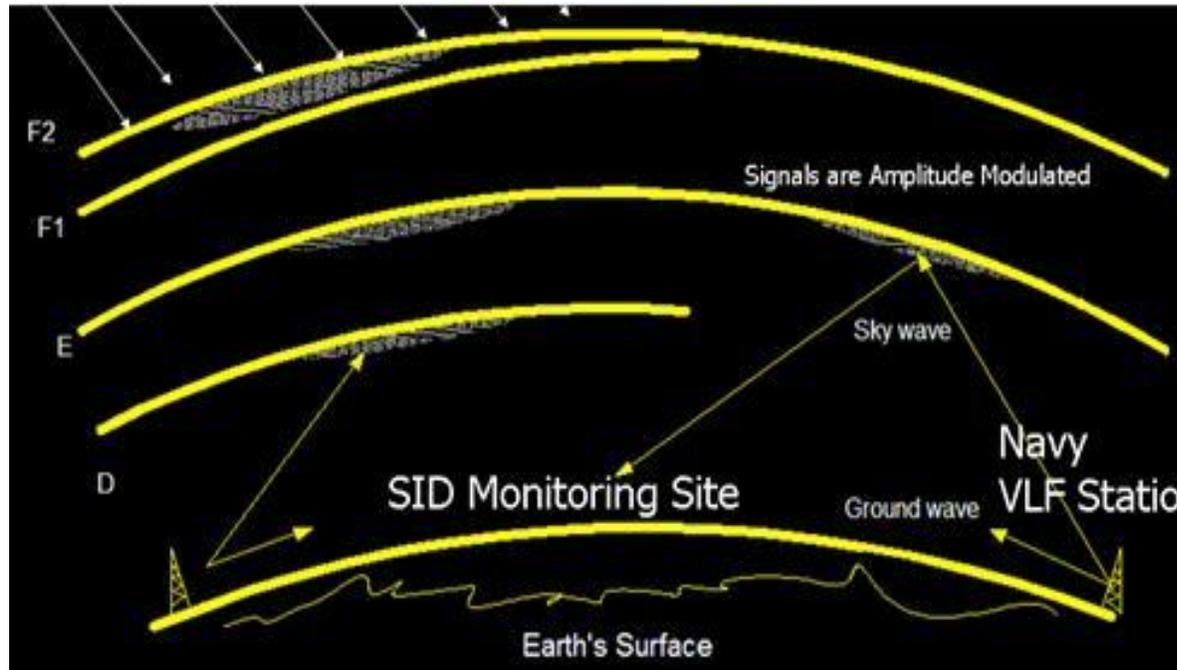
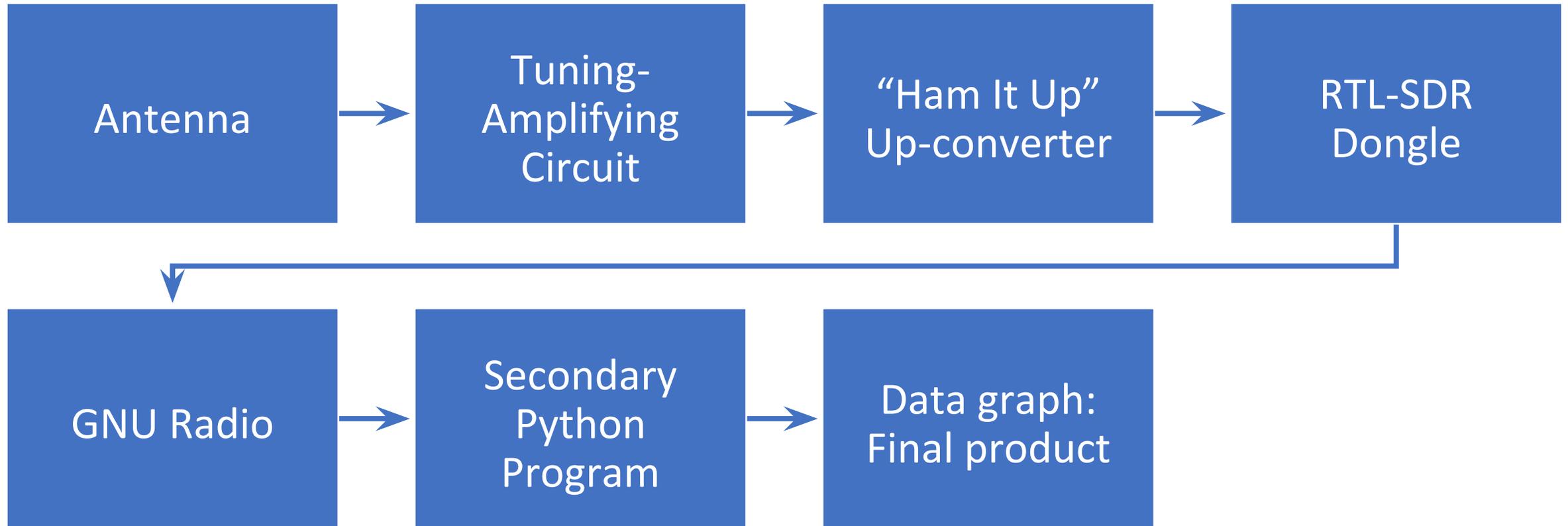
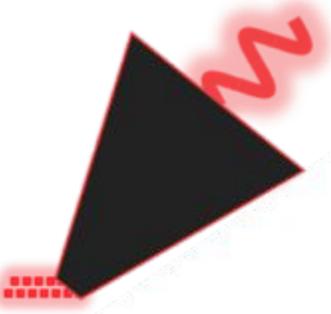


Image source: <http://solar-center.stanford.edu>

- Extra solar activity, like solar flares or solar wind, causes the ionosphere to become more ionized.
- The increase in ionization means that the VLF signal will bounce off the D-Region instead of the higher E or F-Regions, as it normally does
- The signal travels a shorter distance and loses less of its signal strength along the way.
- Therefore, by monitoring the strength of the radio signal, we can detect solar activity!

# Loop Antenna





# Software Comparison: SDR# vs. GNU Radio

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- SDR#:
  - Pros: SDR# is easy to use and requires no setup complications
  - Cons: There is no convenient way to record the signal for further processing, and the DSP steps are less user-adjustable
- GNU Radio:
  - Pros: offers many user-friendly DSP blocks, which can be used to visualize, manipulate and record antenna data
  - Cons: slightly bigger learning curve, displays are not as clear for testing as SDR#'s
- We used SDR# to test the antenna and make sure we were picking up the signals we should be.
- GNU Radio was employed to do the actual DSP and to record the input from the antenna for further processing.

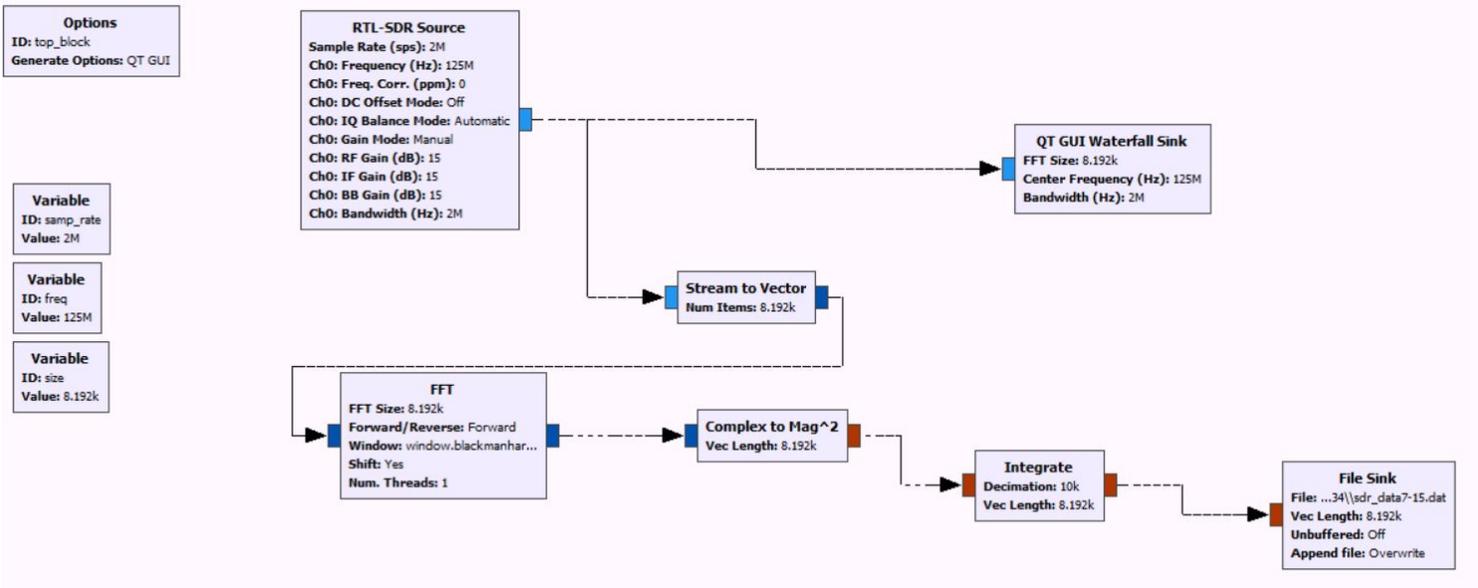


# DSP with GNU Radio

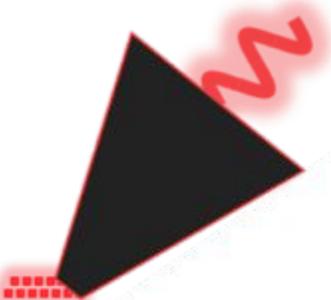
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- SDR dongle as input source
- Set sampling rate to 2M/sec
- Performed 8192 point FFT on the data streamed from the dongle to achieve a spectral resolution of 244 Hz
- Squared the complex FFT output
- Integrated signal every 10k samples
- Saved output to file sink, and displayed in Waterfall Sink

# DSP with GNU Radio

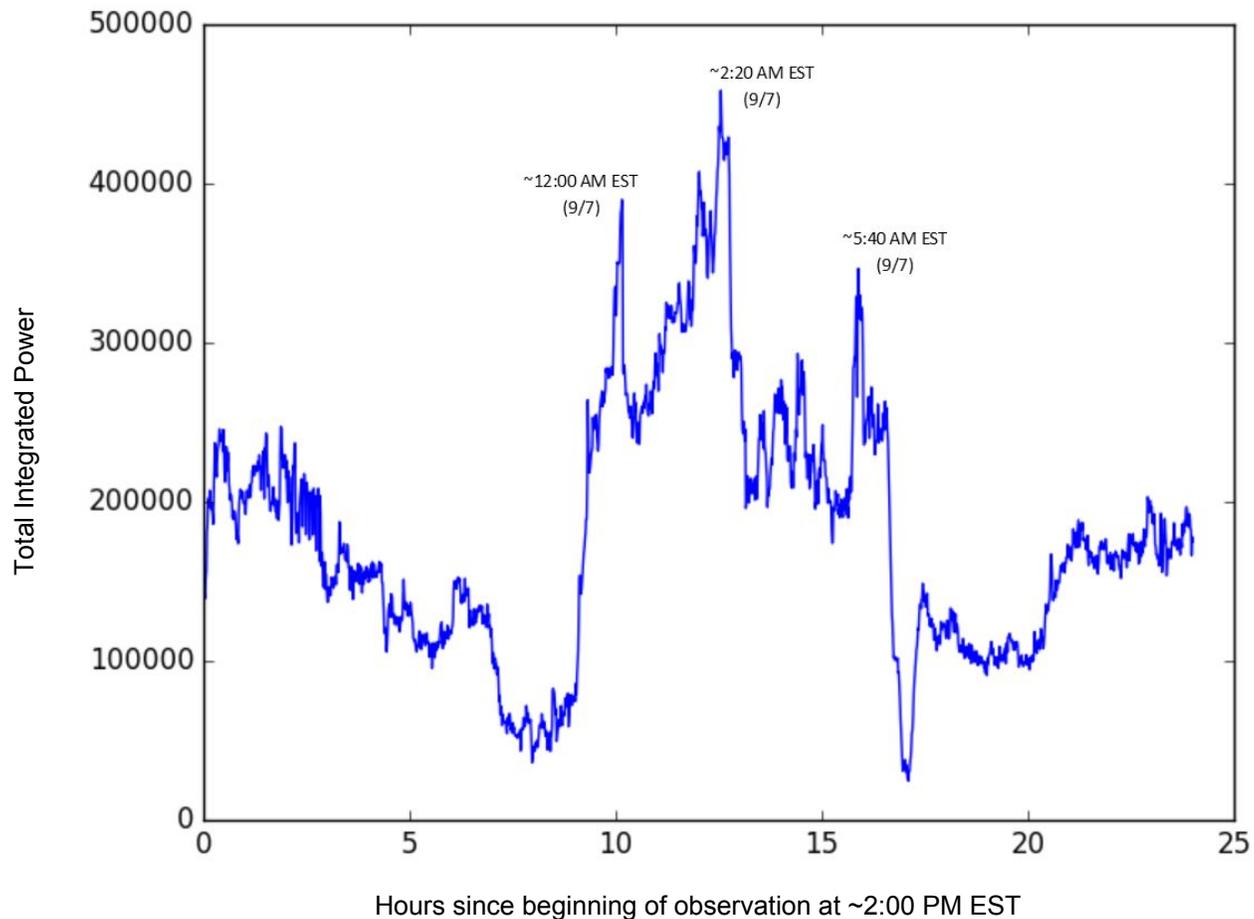


- SDR dongle as input source
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- Performed 8192 point FFT on the data streamed from the dongle to achieve a spectral resolution of 244 Hz
- Squared the complex FFT output
- Integrated signal every 10k samples
- Saved output to file sink, and displayed in Waterfall Sink



# Possible detections...

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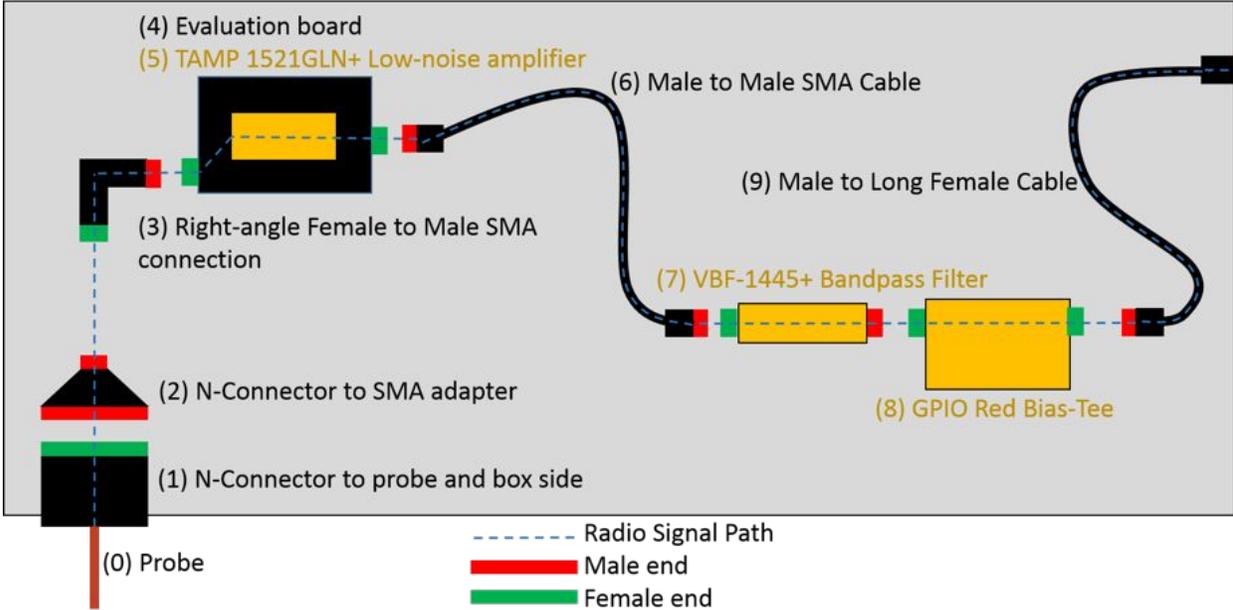


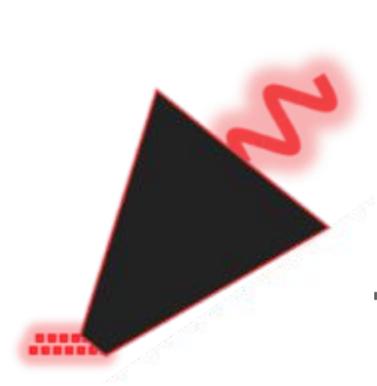
- This data was taken on September 7, 2017, during a time of reported high solar activity.
- The three labelled peaks may be records of solar disturbances in the ionosphere.
- The dips in intensity are likely due to known effects caused by sunrise and sunset.

# The HI Horn Antenna



# Horn Antenna Flowgraph

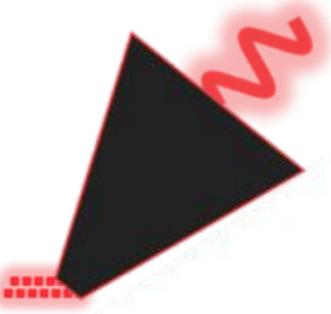




# The idea behind the 21cm horn

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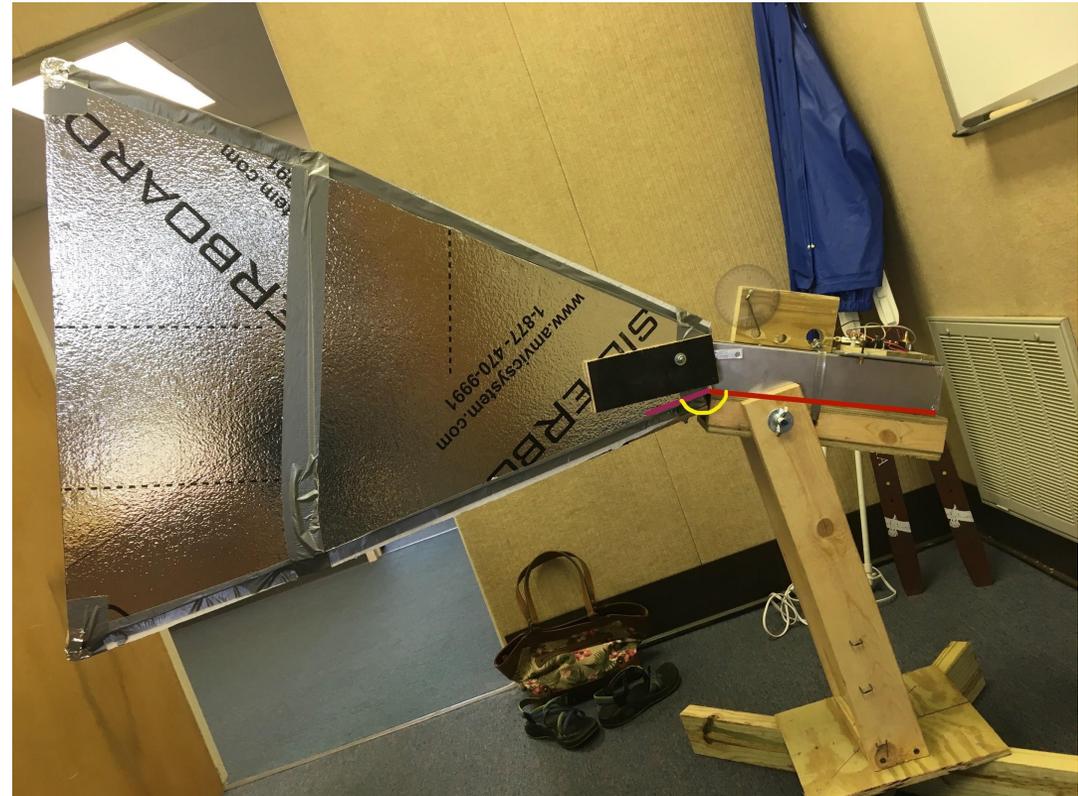
- Mapping the neutral hydrogen of the Milky Way galaxy is possible with an easily-built horn antenna
- The difficulty level of constructing the horn is perfect for high school and undergraduate students
- Provides a great lab/workshop experience
- (~\$250, bit more complicated than loop antenna, about 7 degrees beamwidth)



# Horn construction

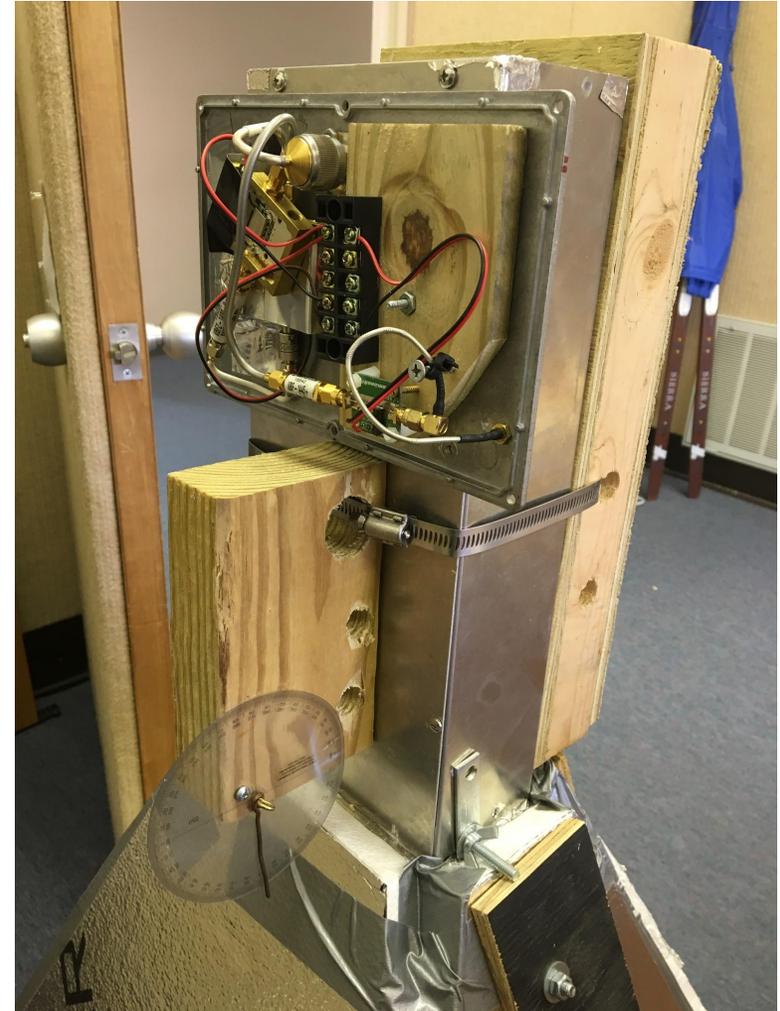
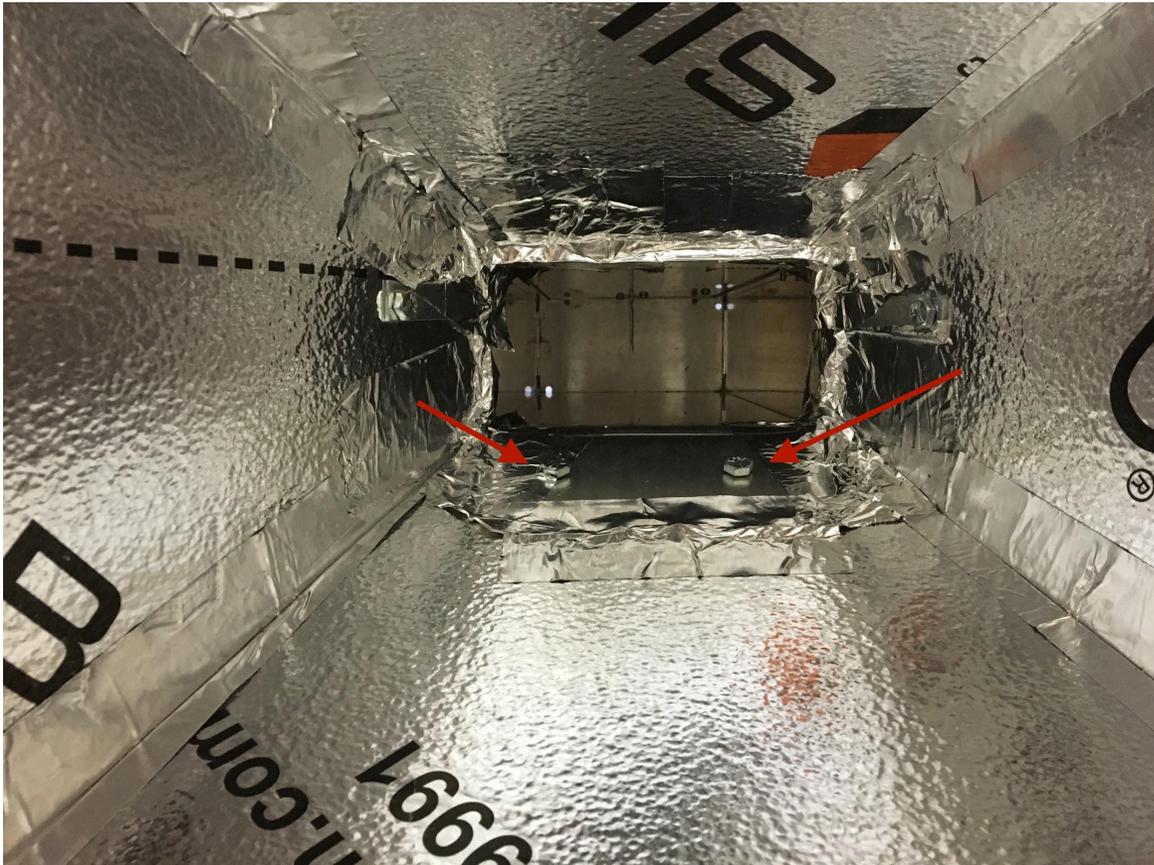
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Horn constructed with styrofoam boards coated with radio-reflective material

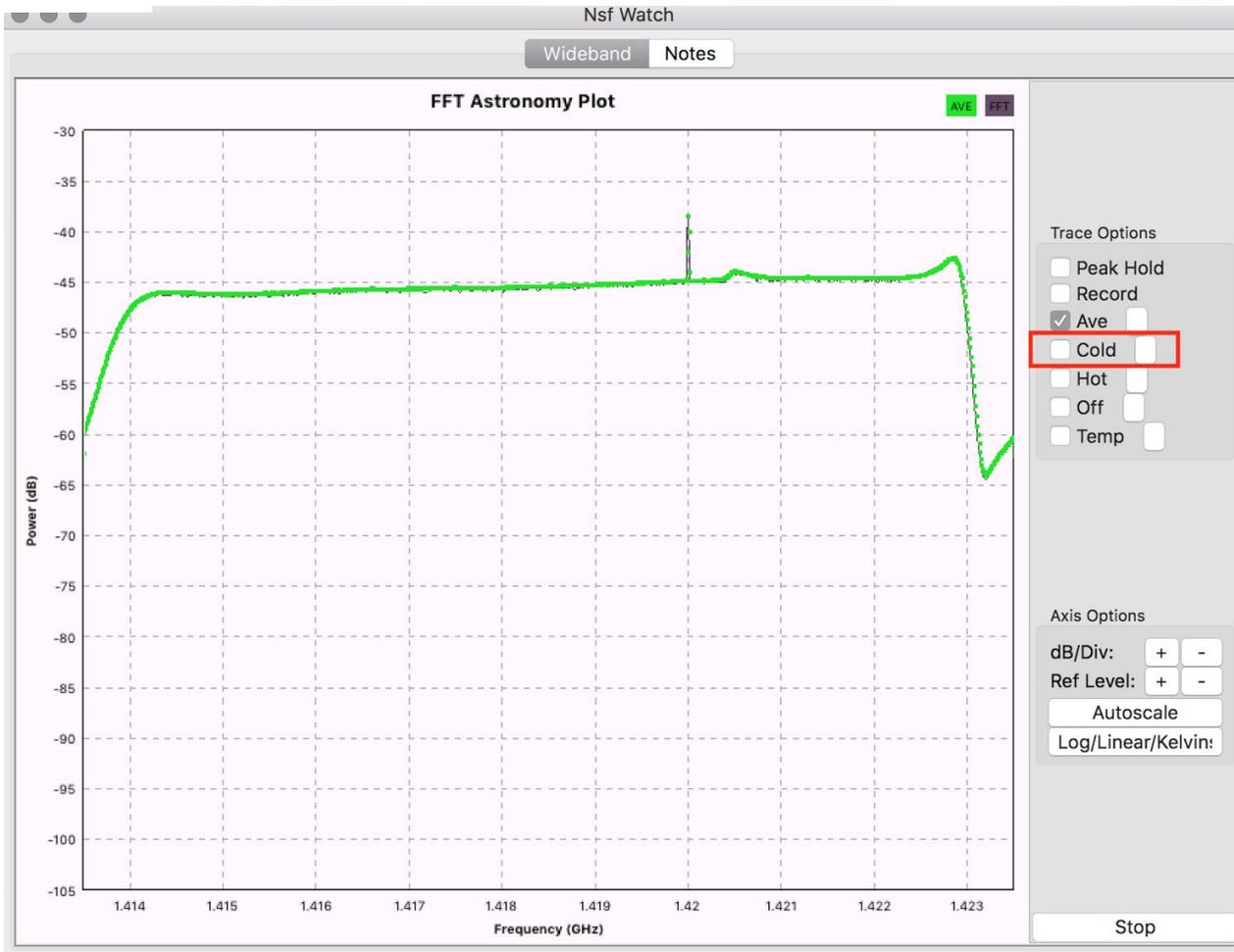


# The Waveguide and LNA

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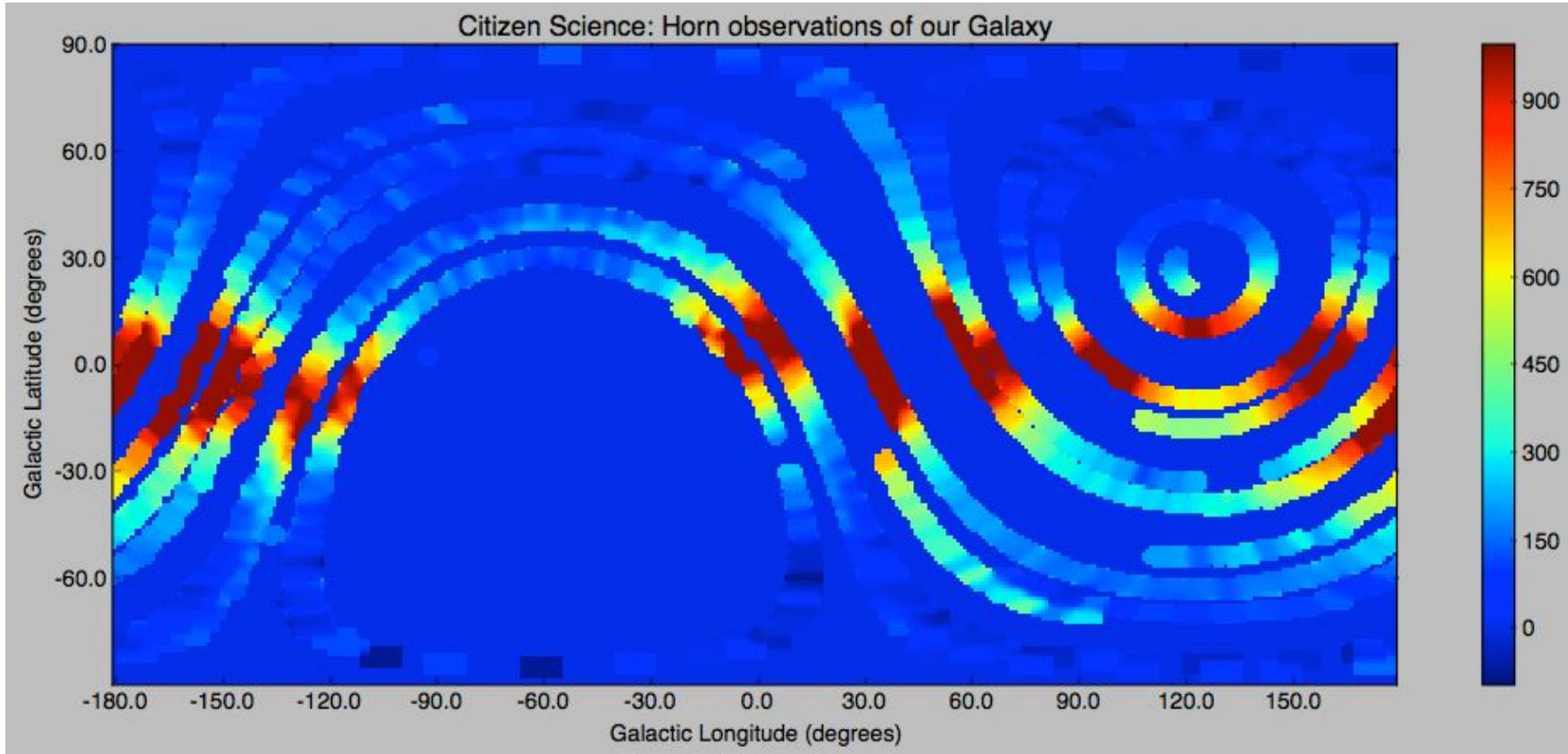


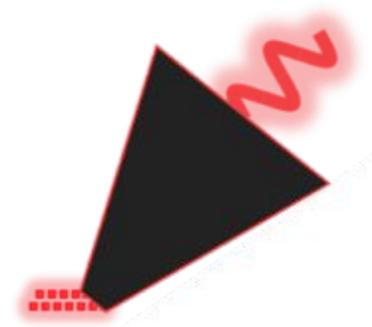
# Digital Signal Processing with GNURadio



```
watch-17apr28 — Python watch.py — 76x34
Setting Azimuth = 180.0
Setting Elevation = 90.0
Setting Sample Rate = 10000000.000000 (Hz)
Setting Integrations= 0:00:10
Gain Names: ('LNA', 'MIX', 'IF')
set_gain 1: 14.0
set_gain 2: 11.0
set_gain 3: 11.0
Setting Obs Note = AficionadoHornTest
New Max number of blocks: 8
msg size: 8192 compared to fft_size 1024
samples in one block 1024, after re-arange 1025
start, 2: 1024,1536 stop, 2: 2048,1537
sample size, sum: 1025,1025:
New Max number of blocks: 9
New Max number of blocks: 10
New Max number of blocks: 13
New Max number of blocks: 14
New Max number of blocks: 15
Not recording! Min: -61.87, Ave: -46.12, Max: -35.84 (db)
Not recording! Min: -62.75, Ave: -46.52, Max: -36.32 (db)
Not recording! Min: -63.18, Ave: -46.73, Max: -36.67 (db)
Not recording! Min: -63.44, Ave: -46.89, Max: -37.76 (db)
Not recording! Min: -63.62, Ave: -46.99, Max: -37.64 (db)
Not recording! Min: -63.77, Ave: -47.04, Max: -37.26 (db)
Not recording! Min: -63.85, Ave: -47.09, Max: -37.20 (db)
Setting Obs Note = AficionadoHorn
Not recording! Min: -63.93, Ave: -47.12, Max: -36.85 (db)
Not recording! Min: -64.00, Ave: -47.17, Max: -38.46 (db)
Setting Integrations= 0:01:00
Setting Azimuth = 0.0
Setting Azimuth = 0.0
Ave: Min 3.98, Max: 1424.58
```

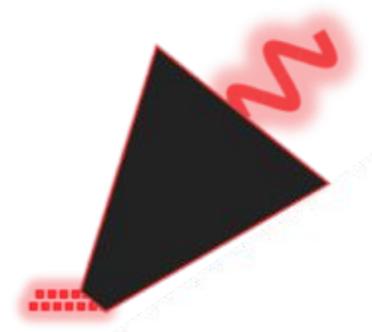
# Observational Results





# How can you get involved with Open Source Radio Telescopes (OSRT)?

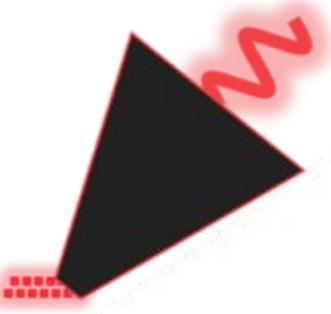
- **Check out our website:** [opensource radiotelescopes.org](https://opensource radiotelescopes.org) and Wiki page, which contain in-progress instructions, educational resources, documentation, and open-source software as the project develops.
- **Teachers:** Build an HI horn or small loop antenna with your students, and share your questions and results with the OSRT community.
- **Join our mailing list!** Visit [opensource radiotelescopes.org/community/](https://opensource radiotelescopes.org/community/)
- You can also **contact us** directly to share your ideas:
  - Ellie White – [orionnebula42@outlook.com](mailto:orionnebula42@outlook.com)
  - Evan Smith – [etsmit12@gmail.com](mailto:etsmit12@gmail.com)
  - Richard Prestage – [richard.prestage@gmail.com](mailto:richard.prestage@gmail.com)



# Summary: Lessons learned

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- Patience and persistence pay off!
- The process of doing a project like this is not as simple as it seems; it takes many attempts and detours to get to a working end result.
- If something works in theory, that doesn't necessarily mean it will work as well in practice.
- The best way to learn about amateur radio astronomy is by doing, and by talking to those who are happy to share their experiences and knowledge.



# Acknowledgements

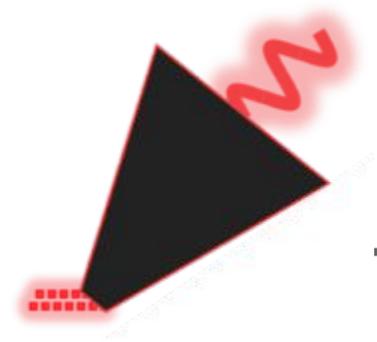
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- Horn design and construction:

- Kevin Bandura
- Sue Ann Heatherly
- Sophie Knudsen
- Glen Langston
- Noreen Prestage
- Richard Prestage
- Pranav Sanghavi
- Evan Smith

- Loop antenna development:

- Ellie White
- Steve White
- Richard Prestage



# Thank you!

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